

Quantum Simulation using Optical Lattices

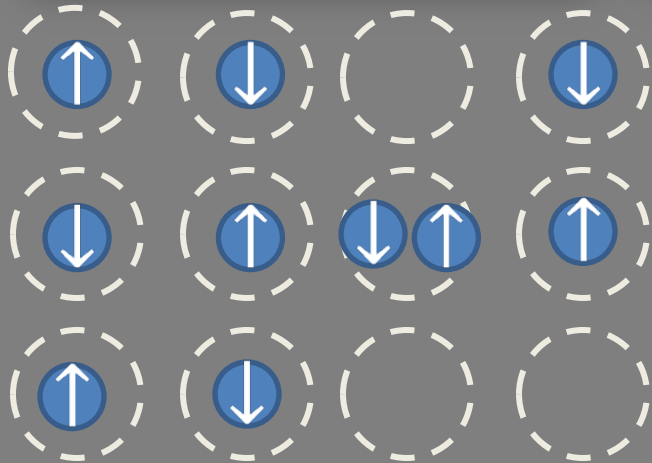
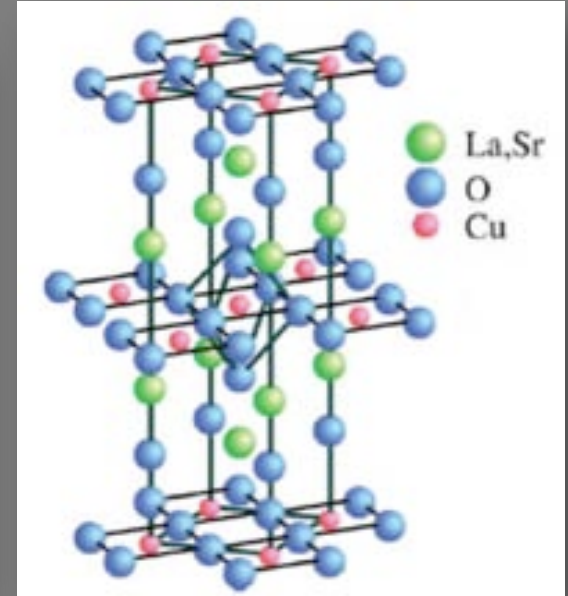
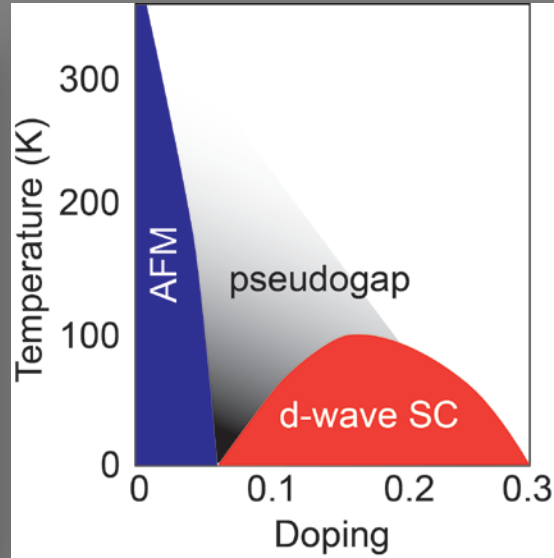
B. DeMarco, University of Illinois at Urbana-Champaign

Agenda:

- Basic Physics of Atoms in Optical Lattices
- Thermometry and Cooling; Spin-Dependent Lattices
- Disordered Gases

The Premise: an example

High temperature superconductivity / Hubbard model



$$H = -t \sum_{\langle ij \rangle, \sigma} c_{i, \sigma}^{\dagger} c_{j, \sigma} + U \sum_i n_{i, \uparrow} n_{i, \downarrow}$$

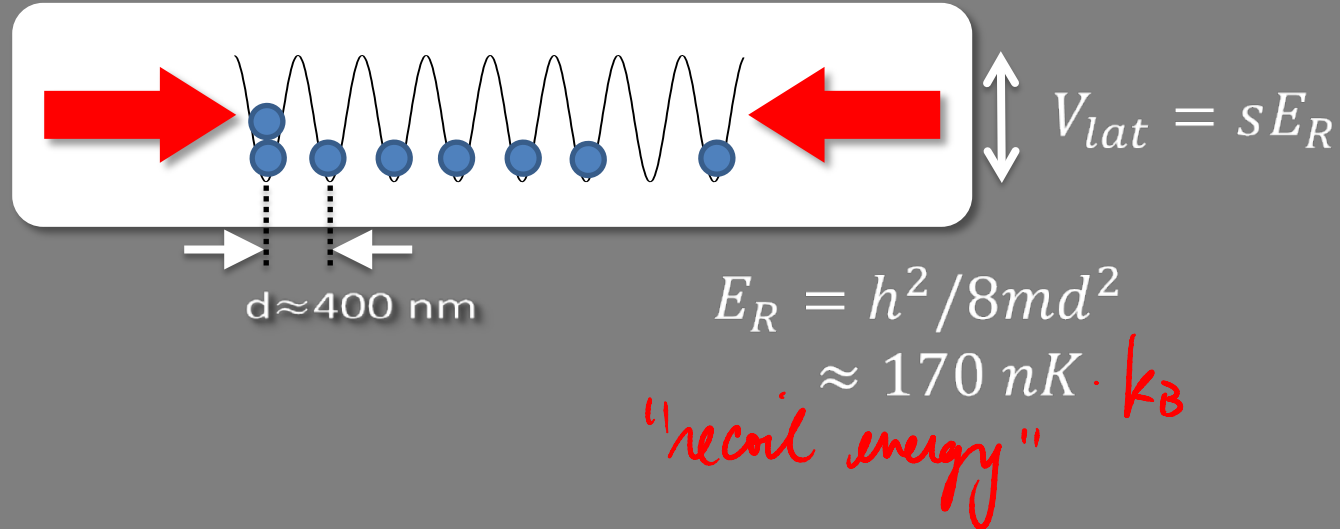
tunneling

interactions

- Can Hubbard model produce d-wave SC?
- Nature of pseudo-gap phase?

Optical Lattice Quantum Simulation

Atoms confined by periodic potential arising from intensity and/or polarization gradients



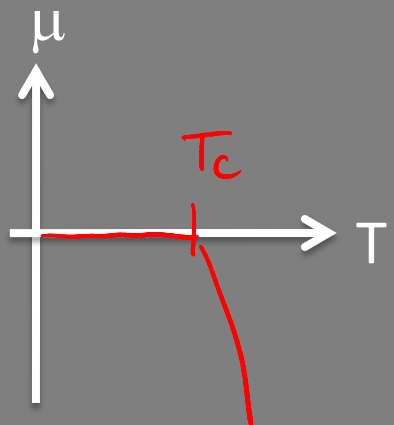
“Exactly” realizes the Hubbard model

Today:

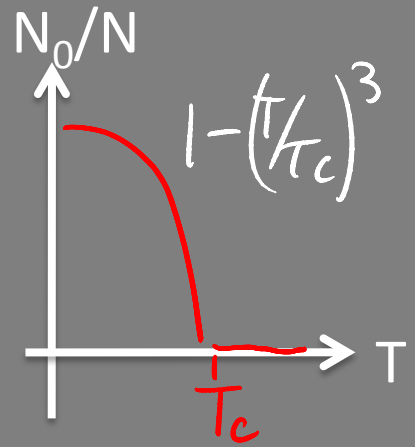
- A on little cooling / quantum degeneracy
- Optical lattice potentials / AC Stark Effect
- Interactions (collisions)

Quantum Degeneracy (for an ideal gas)

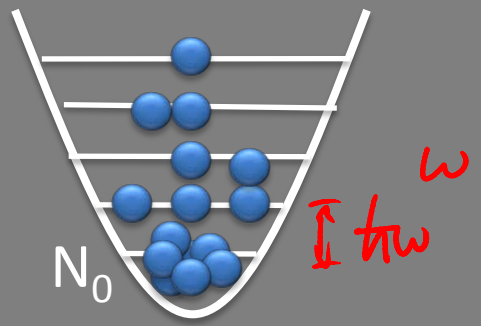
Bosons



$$T_c = 0.94 \hbar \omega N^{1/3}$$

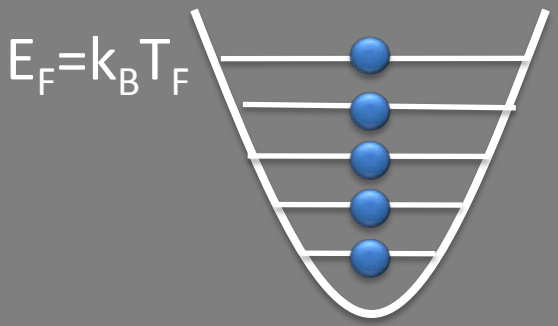
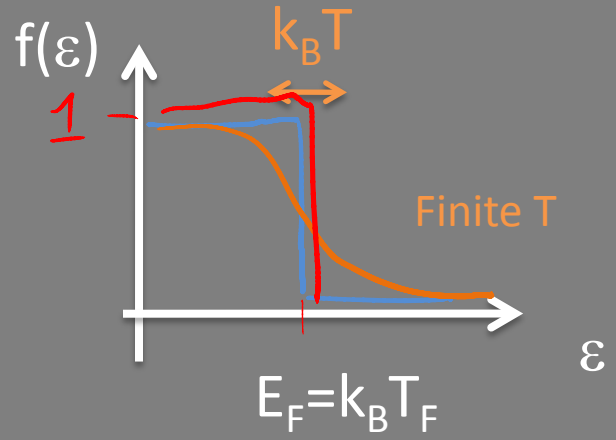


$$n_{pk} \Lambda_{dB}^3 = 2.6$$

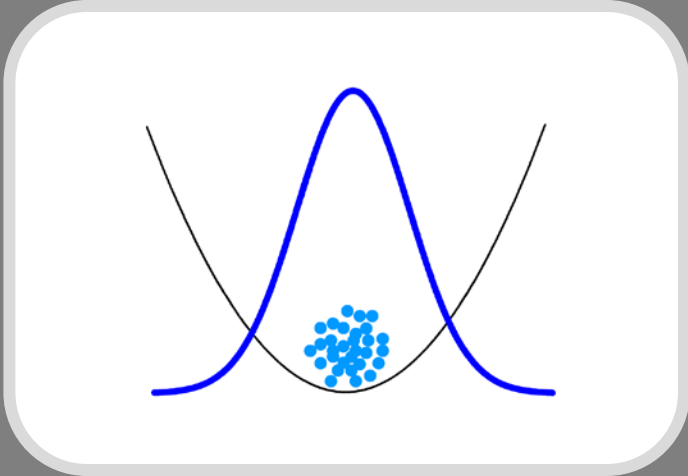


Fermions

$$E_F = \hbar \omega (6N)^{1/3}$$

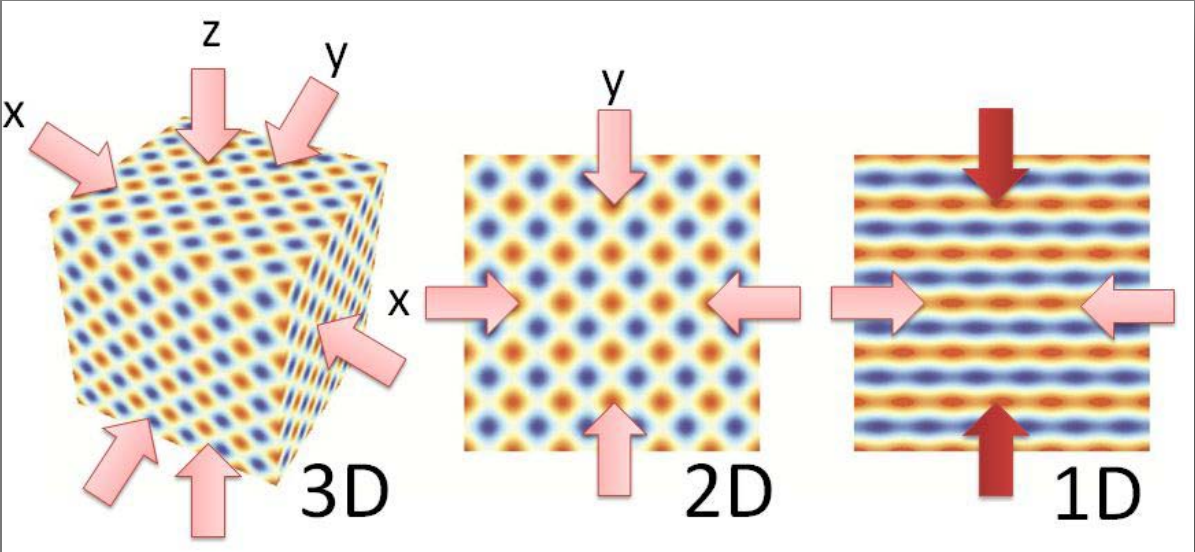


Lattice Potentials



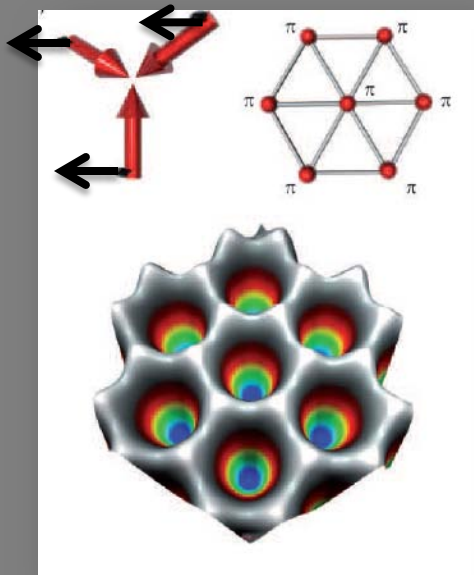
Superimpose lattice potential on trapped, degenerate gas

Simplest geometry: cubic, square, 1D



Other geometries

triangular



hexagonal

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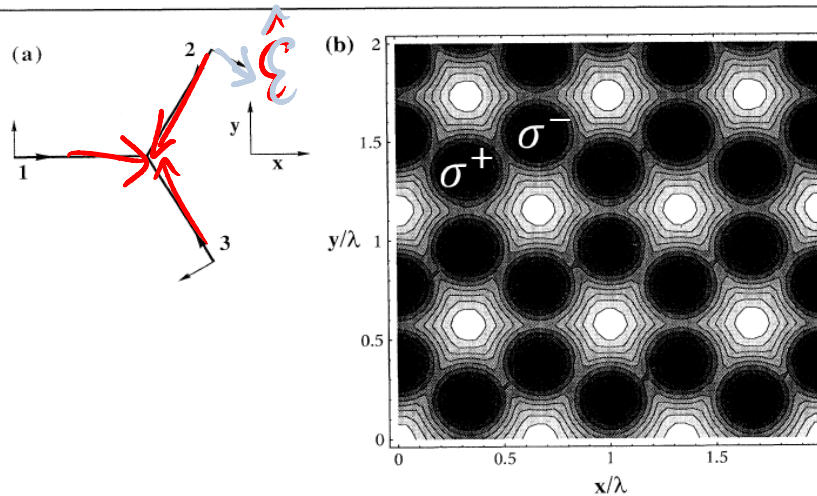


FIG. 1. Beam configuration for the 2D molasses. (a) Three coplanar beams of equal intensity have wave vectors making a 120° angle with each other. The three beams are linearly polarized in the plane of the figure. (b) Spatial dependence of the minima of the optical potential. The scale on both the x and y axes is the laser wavelength λ . The potential wells have their minima (which appear in black) on a hexagonal lattice, at points where the light polarization is purely circular.

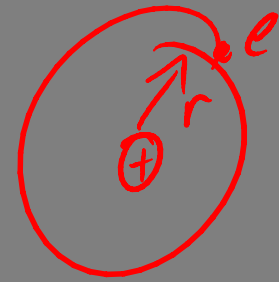
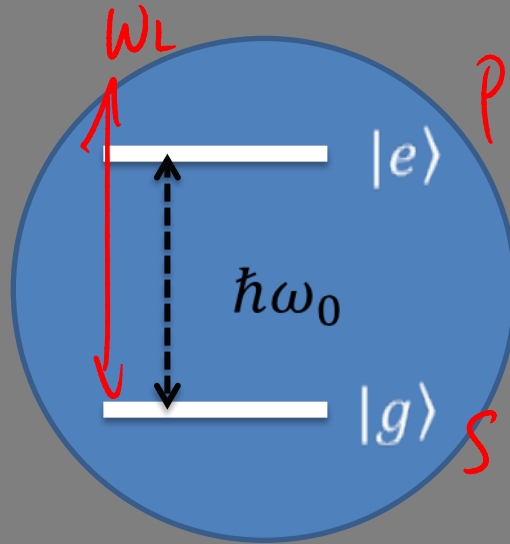
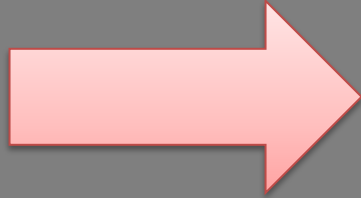
Sengstock

Spin-dependent: different F, m_F experience different potentials
(more tomorrow)

AC Stark Effect

$$\vec{E} = E_0 \hat{e} \cos(kz - \omega_L t)$$

↑
polarization



$$H_I = e \vec{r} \cdot \vec{E} \cos(\omega_L t)$$

dipole matrix element

Treat atom as 2-level system
(electric dipole approximation)

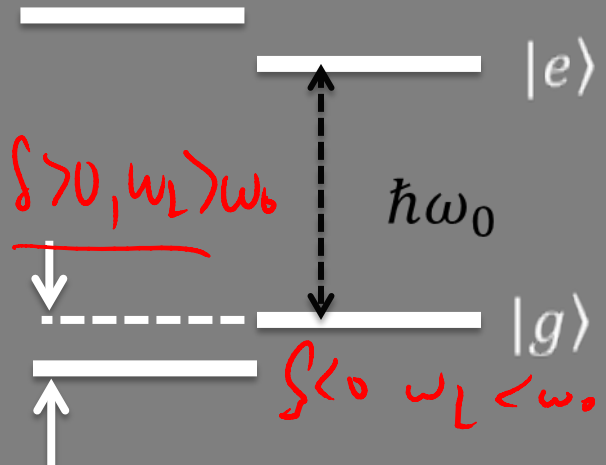
$$\text{Rabi frequency: } \Omega = \frac{eE_0}{2} \langle e | \vec{r} \cdot \hat{e} | g \rangle$$

Rotating wave approximation,
 $\Omega \ll \delta = \omega_L - \omega_0$

$$\Delta E \propto I / \delta$$

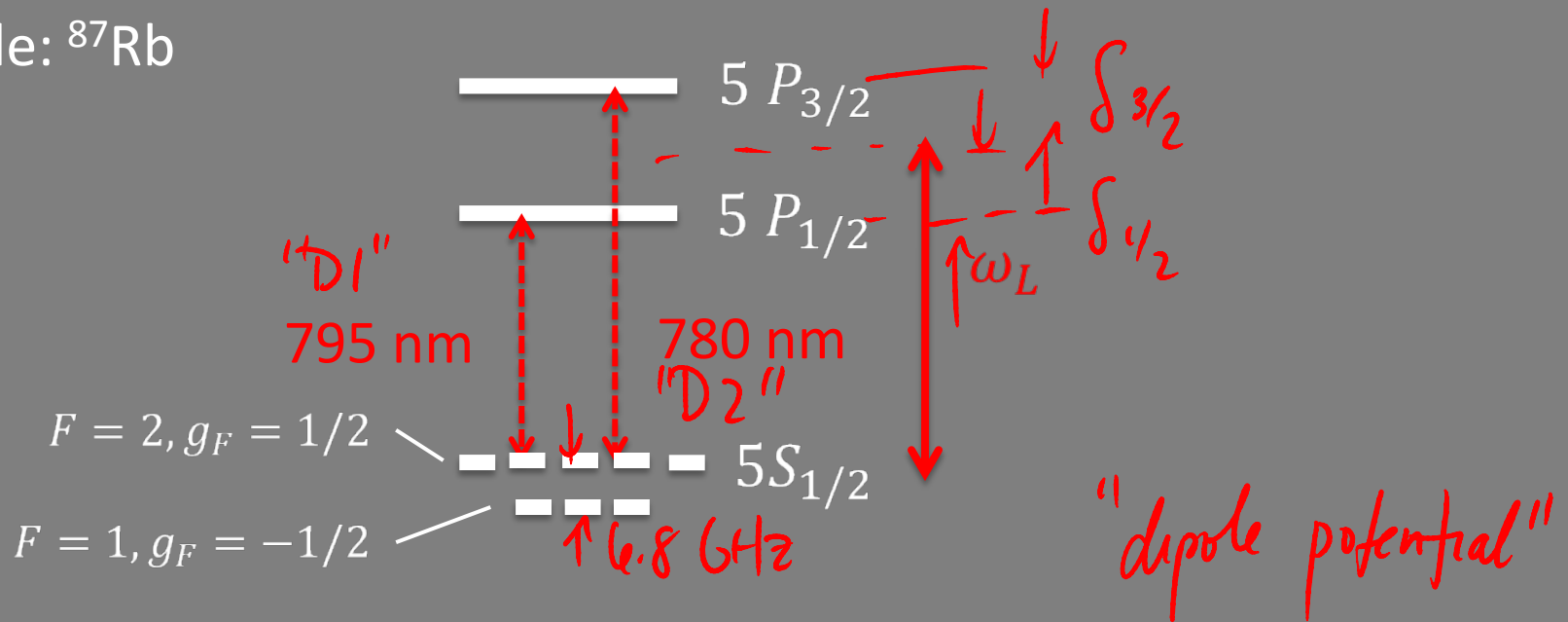
$$I \propto |E_0|^2$$

$$\hbar |\Omega|^2 / \delta$$



Real Atoms

Example: ^{87}Rb



$$U_{dip} = \frac{\pi c^2 \Gamma}{2\omega_0} \left(\frac{2 + P g_F m_F}{\delta_{3/2}} + \frac{1 - P g_F m_F}{\delta_{1/2}} \right) \frac{I(\vec{r})}{|\Omega|^2 \propto |E_0|^2}$$

$P = -1 (\sigma^-), 0 (\pi), 1 (\sigma^+)$ Polarization *in the atomic basis*

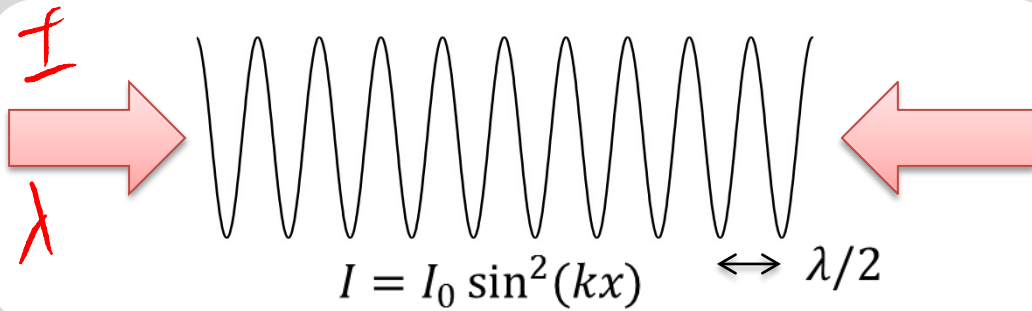
$$\Gamma = \frac{\omega_0^3}{3\pi\epsilon_0\hbar c^3} |\langle e | e\vec{r} | g \rangle|^2 = 1/\tau \approx 2\pi \times 6\text{ MHz}$$

Spin-Independent Lattice

π -polarized light (or linear in any basis), or large detunings

1064 nm

$$U_{dip} = \frac{3\pi c^2 \Gamma}{2\omega_0 \delta} I(\vec{r})$$



$U \propto \sin^2(kx)$

Red detuning:

$\delta < 0$, trapped in anti-nodes

Blue detunings:

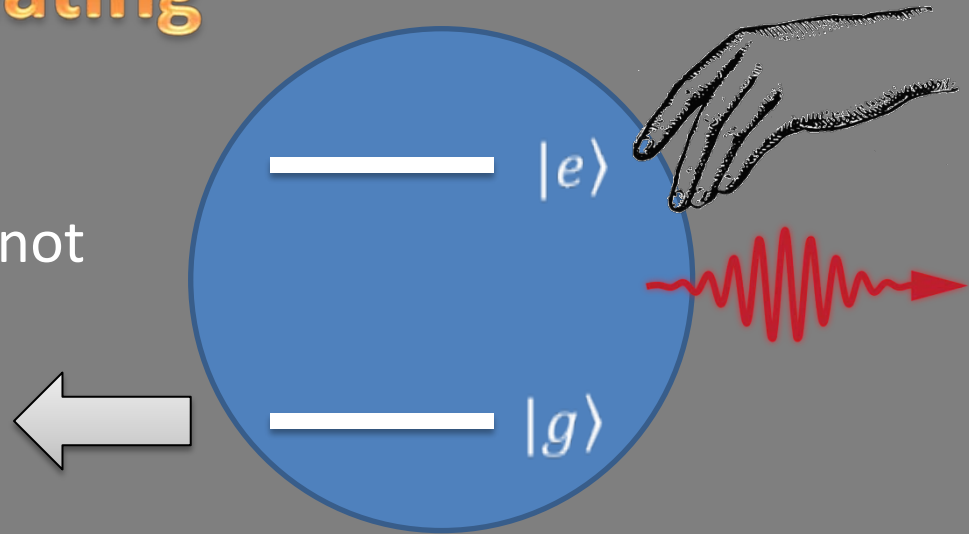
$\delta > 0$, trapped in nodes

Note corrections for large detunings
($\lambda = 532, 1064, 1550 \text{ nm}$)

$$U_{dip} = -\frac{3\pi c^2 \Gamma}{2\omega_0} \left(\frac{1}{\omega_0 - \omega_L} + \frac{1}{\omega_0 + \omega_L} \right) I(\vec{r})$$

Light-induced Heating

Electromagnetic vacuum is not our friend (sometimes)



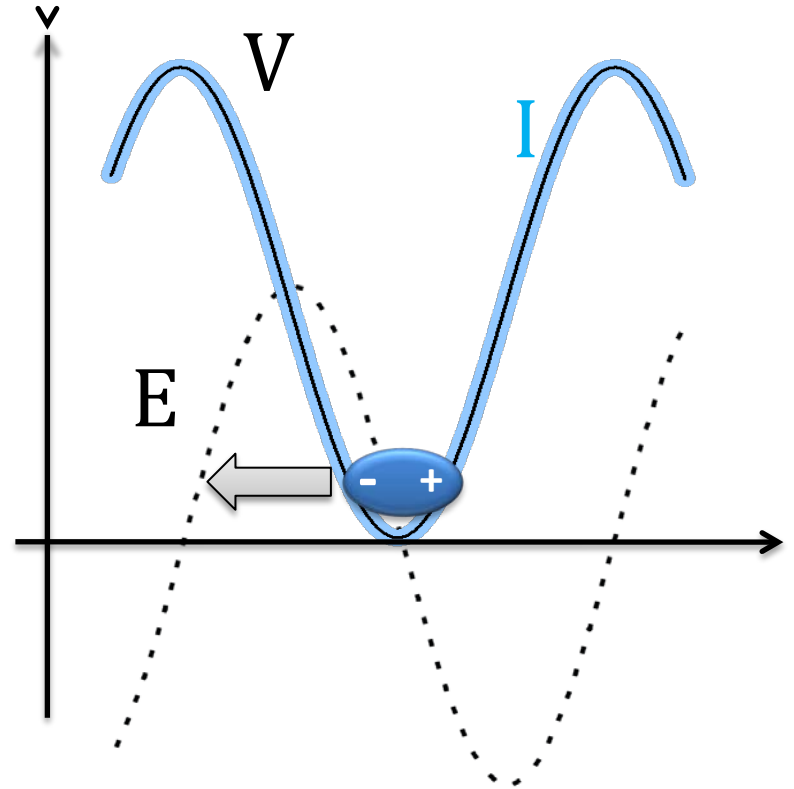
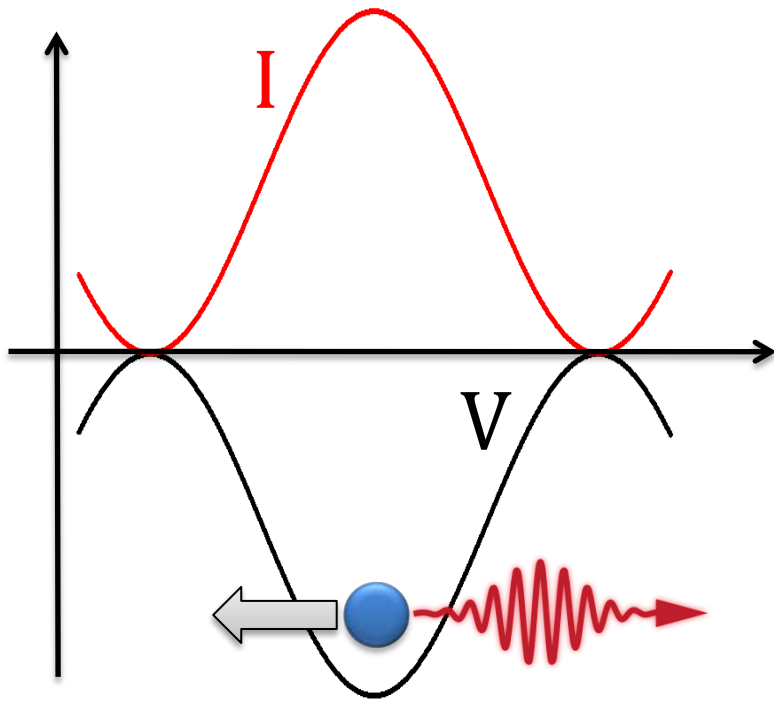
So, it's better to keep the atoms in the dark in a blue-detuned lattice, right?

Gordon and Ashkin: it doesn't matter

I_0

$\dot{E} = \frac{E_R^3 6\pi c^2 \Gamma^2}{\hbar \omega_0^3 \delta^2} I_0 \left[\left(\frac{\omega}{\omega_0} \right)^2 \frac{E_R^3 6\pi c^2 \Gamma^2}{\hbar \omega_0^3 \delta^2} \frac{I_0}{(\omega_0 - \omega_L)^2} + \frac{1}{(\omega_0 + \omega_L)^2} \right]$

Light Induced Heating



Result the same for spin-dependent lattices,
imbalanced beams, beams at an angle...

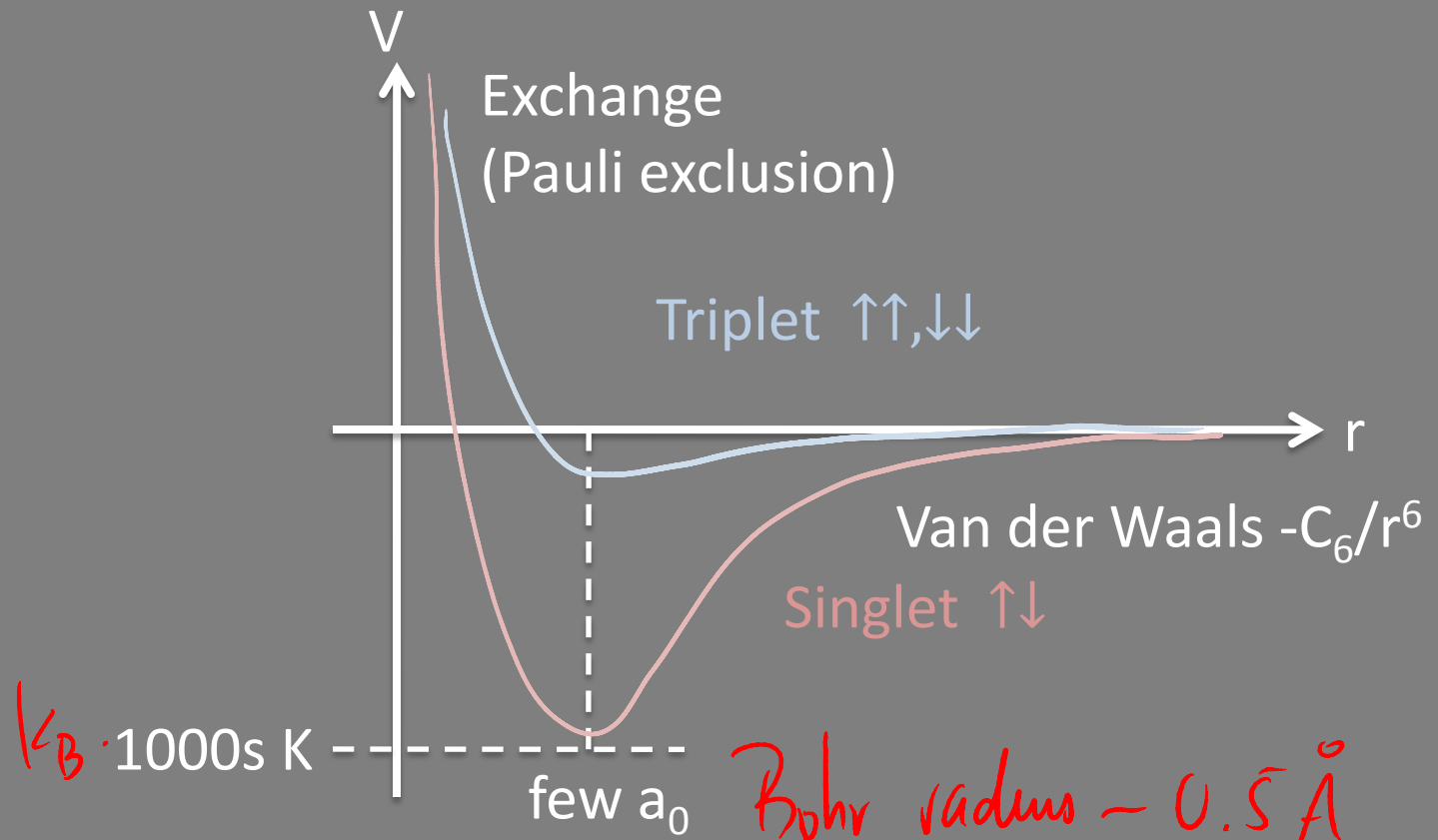
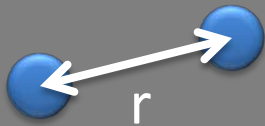
Ultra-cold collisions

$$H = -t \sum_{\langle ij \rangle, \sigma} c_{i, \sigma}^\dagger c_{j, \sigma} + U \sum_i n_{i, \uparrow} n_{i, \downarrow}$$

Where does U come from?

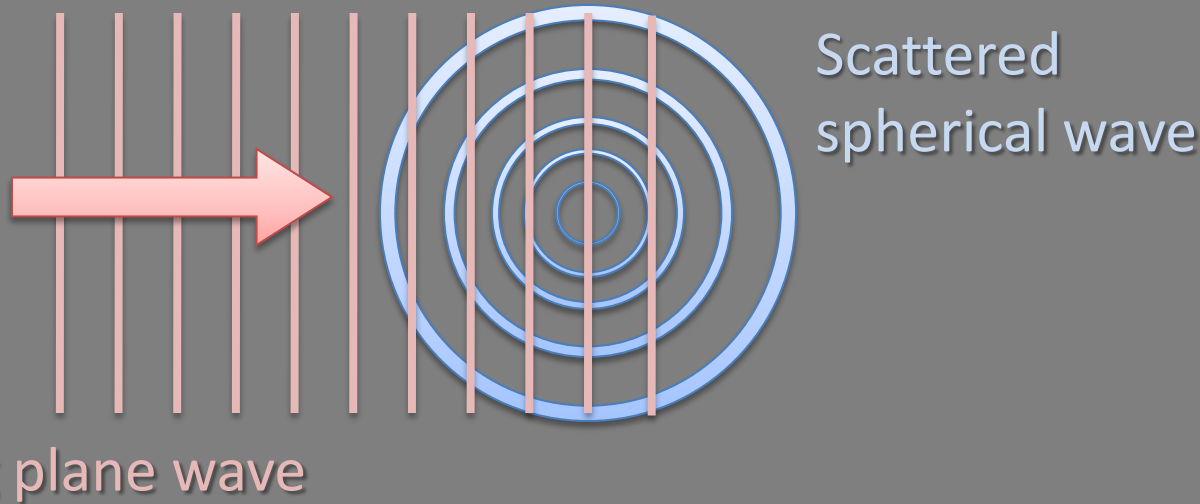
a: scattering length

One number that characterizes binary, low-energy collisions!



Ultra-cold collisions

Very basic quantum scattering theory



Goal: find the scattered wave

Lippman-Schwinger

$$\Psi \propto e^{ikz} + f(k, \Omega) \frac{e^{ikr}}{r} \quad f: \text{scattering amplitude}$$

$$\Psi = \sum_l R_l(k, r) P_l(\cos \vartheta) \quad \left[\frac{\partial^2}{\partial r^2} + \frac{2}{r} \frac{\partial}{\partial r} - \frac{l(l+1)}{r^2} - V(r) + k^2 \right] R_l(k, r) = 0$$

partial wave expansion

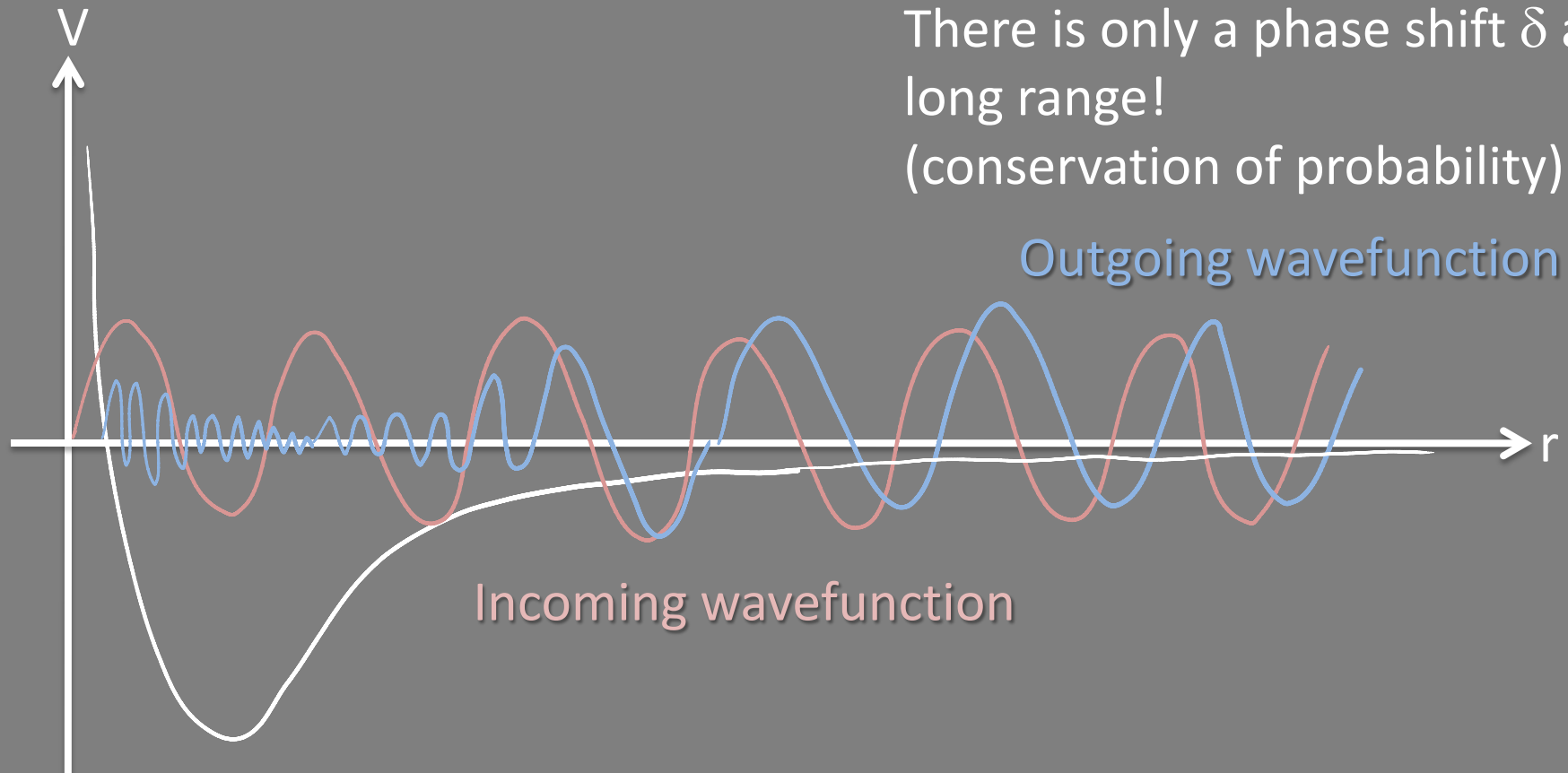
centrifugal barrier

Ultra-cold collisions

Very basic quantum scattering theory

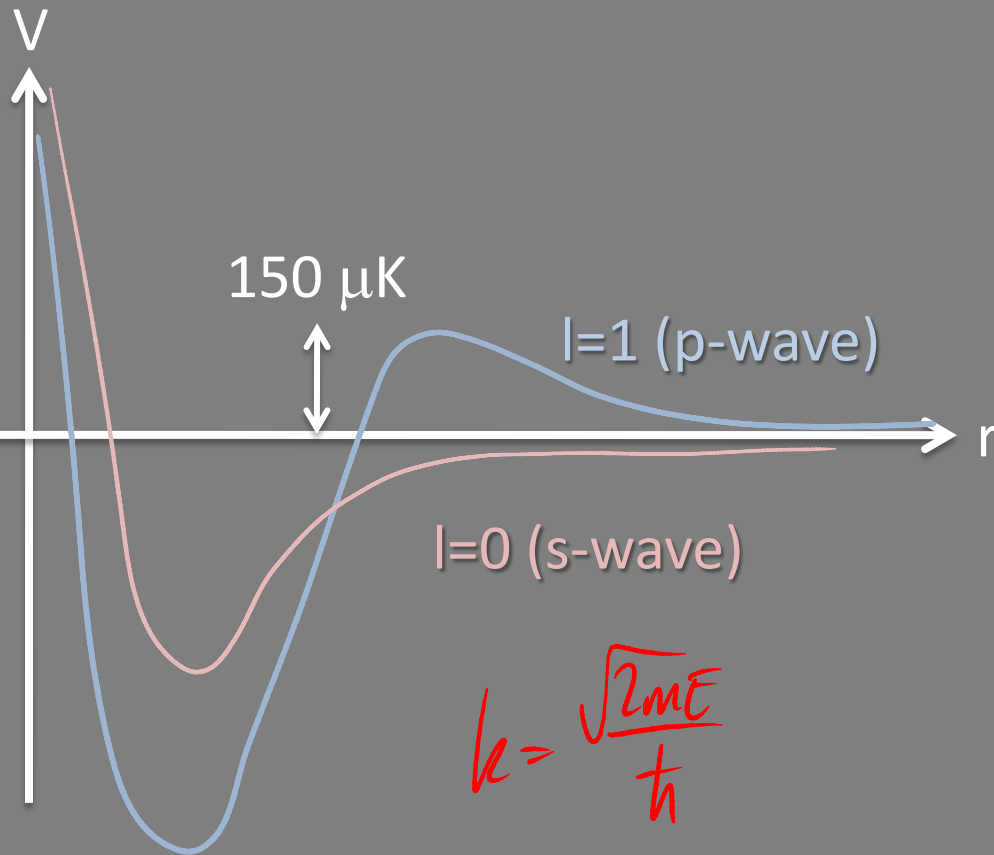
Asymptotic solution $R_l \underset{r \rightarrow \infty}{\propto} \frac{1}{kr} \sin \left[kr - l \frac{\pi}{2} + \delta_l(k) \right]$

There is only a phase shift δ at long range!
(conservation of probability)



Ultra-cold collisions

Very basic quantum scattering theory



At ultra-cold temperatures, we only have to consider s-wave ($l=0$) collisions

$$R_0 \underset{r \rightarrow \infty}{\propto} \sin(kr + \delta_0) / kr$$

$$R_0 \underset{r \rightarrow \infty}{\propto} \sin[k(r - a)] / kr$$

Trade phase shift for a length

$$a = -\lim_{k \rightarrow 0} \frac{\delta_0(k)}{k}$$

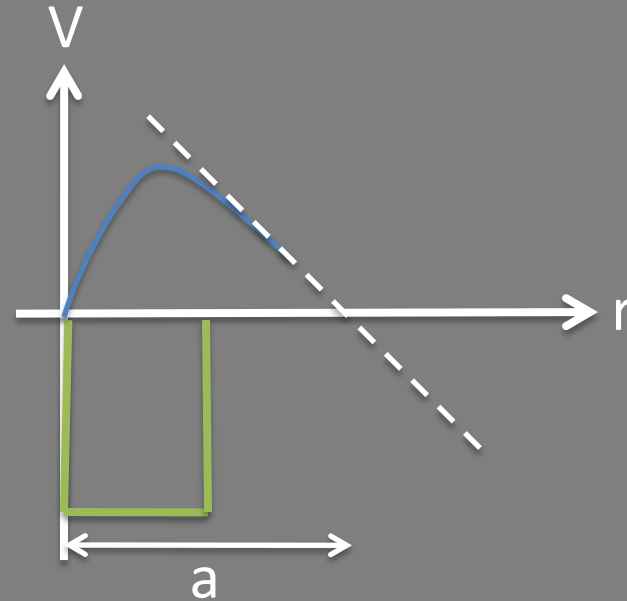
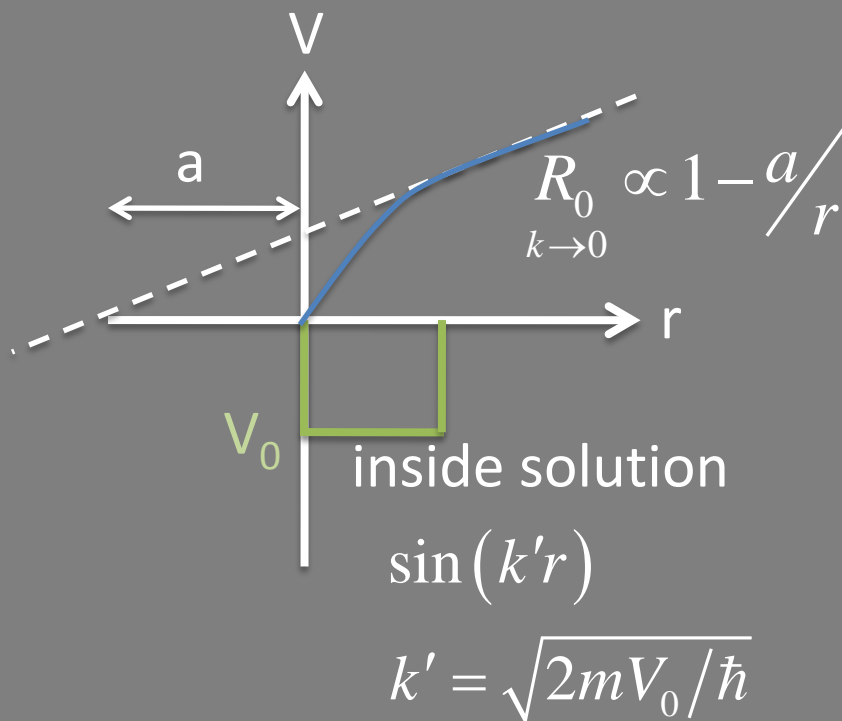
$$a = -\infty \rightarrow \infty$$

Ultra-cold collisions

Very basic quantum scattering theory

Examples: ^{133}Cs $a = -3000 a_0$ ^{40}K $a = 200 a_0$ (at $B=0$)

How can a take on any value?



$a > 0$ or < 0 related to last bound state energy
...Feshbach resonances on Friday

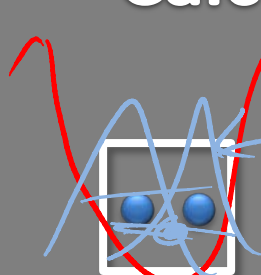
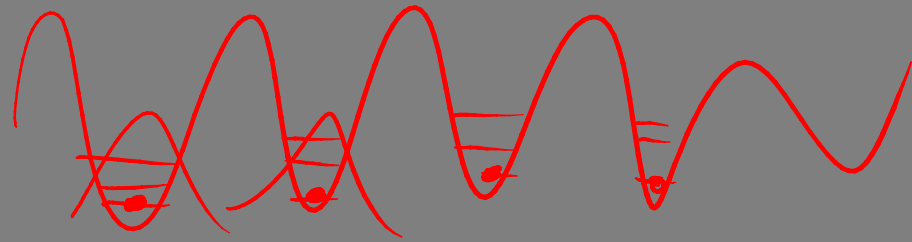
Contact interaction approximation

Interactions treated using a pseudo-potential (see Huang)

$$V(r) = \frac{4\pi\hbar^2}{m} a \delta^3(\vec{r})$$

- Long range behavior correct $R \propto 1 - a/r$
- Enforces boundary condition $\Psi(r=a) = 0$

Calculating U



$$E_{\text{int}} \propto \int d^3\vec{x}_1 \int d^3\vec{x}_2 \phi_0^*(\vec{x}_1) \phi_0^*(\vec{x}_2) \frac{4\pi\hbar^2 a}{m} \delta^3(\vec{x}_1 - \vec{x}_2) \phi_0(\vec{x}_1) \phi_0(\vec{x}_2)$$

$$U \propto \frac{4\pi\hbar^2 a}{m} \int d^3\vec{x} |\phi_0(\vec{x}_1)|^4$$

Ultra-cold collisions

Very basic quantum scattering theory

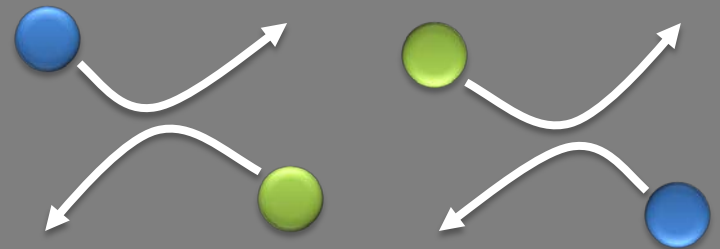
bosons vs fermions

$$\Psi = \Psi_{space} \left| \chi_{spin} \right\rangle \propto P_l(\cos \mathcal{G}) \left| \chi_{spin} \right\rangle$$

Two-particle wavefunction must have proper exchange symmetry

$$\Psi(x_1, x_2) \rightarrow \Psi(x_2, x_1) \quad \text{bosons}$$

$$\Psi(x_1, x_2) \rightarrow -\Psi(x_2, x_1) \quad \text{fermions}$$



equivalent to 180° rotation

$$P_l(\cos \mathcal{G}) = P_l(-\cos \mathcal{G}) \quad l \text{ even}$$

$$P_l(\cos \mathcal{G}) = -P_l(-\cos \mathcal{G}) \quad l \text{ odd}$$

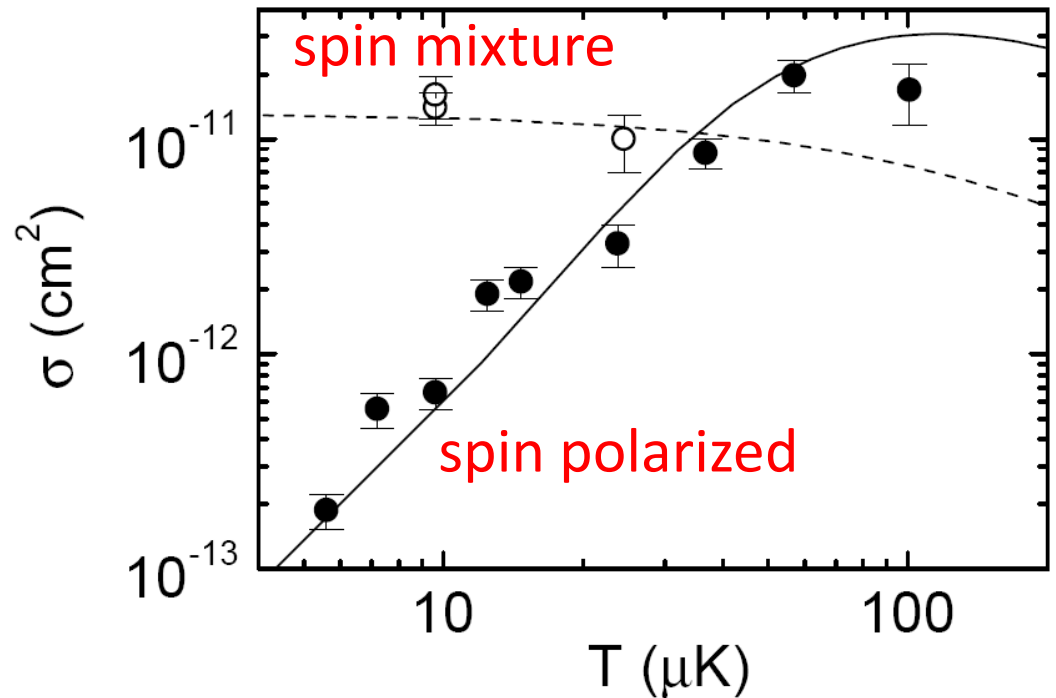
Ultra-cold collisions

Very basic quantum scattering theory

At low temperatures, $l=0$: the spatial wavefunction must be symmetric

$|\chi\rangle = (|\uparrow\downarrow\rangle - |\downarrow\uparrow\rangle) / \sqrt{2}$ Fermions: require spinor to be anti-symmetric

Multiple spin states are required for fermionic atoms to interact!



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